# Trace formula for noise corrections to trace formulas 

Gergely Palla, ${ }^{1}$ Gábor Vattay ${ }^{1}$, and André Voros ${ }^{2}$<br>${ }^{1}$ Department of Physics of Complex Systems, Eötvös University, Pázmány Péter sétany 1/A, H-1117 Budapest, Hungary<br>${ }^{2}$ CEA, Service de Physique Théorique de Saclay F-91191 Gif-sur-Yvette CEDEX, France

(Received 3 October 2000; published 15 June 2001)


#### Abstract

We consider an evolution operator for a discrete Langevin equation with a strongly hyperbolic classical dynamics and Gaussian noise. Using an integral representation of the evolution operator $\mathcal{L}$, we investigate the high-order corrections to the trace of $\mathcal{L}^{n}$. The asymptotic behavior is found to be controlled by subdominant saddle points previously neglected in the perturbative expansion. We show that a trace formula can be derived to describe the high-order noise corrections.


DOI: 10.1103/PhysRevE.64.012104
PACS number(s): 05.40.-a, 05.45.-a, 03.65.Sq

In the statistical theory of dynamical systems the development of the densities of particles is governed by a corresponding evolution operator. For a repeller, the leading eigenvalue of this operator $\mathcal{L}$ yields a physically measurable property of the dynamical system, the escape rate from the repeller. In the case of deterministic flows, the periodic orbit theory [1,2] yields explicit and numerically efficient formulas for the spectrum of $\mathcal{L}$ as zeros of its spectral determinant [3].

Stochastic processes of various strength have an influence on all dynamical evolutions. In a series of papers [4-7] the effects of noise on measurable properties such as dynamical averages in classical chaotic dynamical systems were systematically accounted for. The theory developed is closely related to the semiclassical $\hbar$ expansions [8-10] based on Gutzwiller's formula for the trace in terms of classical periodic orbits [11] in that both are perturbative theories in the noise strength or $\hbar$, derived from saddle-point expansions of a path integral containing a dense set of unstable stationary points. The analogy with quantum mechanics and field theory is made explicit in [4] where Feynman diagrams are used to find the lowest nontrivial noise corrections to the escape rate.

In [6] we developed an explicit matrix representation of the stochastic evolution operator. The numerical implementation made it possible to reach up to order eight in expansion order, and the corrections to the escape rate were found to be a divergent series in the noise expansion parameter. This reflects that the corrections were calculated (using the so-called cumulant expansion) from other divergent quantities, the traces of the evolution operator $\mathcal{L}^{n}$ [6].

In [7] the focus was on the high-order noise corrections for the special case of the first trace, $\operatorname{Tr} \mathcal{L}$. The asymptotics of the trace of the evolution operator were governed by subdominant saddles previously neglected in the expansion.

In this paper we show that the high-order noise corrections of $\operatorname{Tr} \mathcal{L}^{n}$ are also dominated by subdominant saddles. These subdominant saddles can be treated as generalized periodic orbits of the system and we associate them with periodic orbits of corresponding discrete Newtonian equations of motion. Our key result is Eq. (40) where the high-order noise corrections are converted into a trace formula. We give as a numerical example the quartic map considered in [4-7].

First we introduce the noisy repeller and its evolution operator. An individual trajectory in the presence of additive noise is generated by iterating

$$
\begin{equation*}
x_{n+1}=f\left(x_{n}\right)+\sigma \xi_{n}, \tag{1}
\end{equation*}
$$

where $f(x)$ is a map, $\xi_{n}$ a random variable with the normalized distribution $p(\xi)$, and $\sigma$ parametrizes the noise strength. In what follows we shall assume that the mapping $f(x)$ is one-dimensional and expanding, and that the $\xi_{n}$ are uncorrelated. A density of trajectories $\phi(x)$ evolves with time on the average as

$$
\begin{equation*}
\phi_{n+1}(y)=\left(\mathcal{L}^{\circ} \phi_{n}\right)(y)=\int d x \mathcal{L}(y, x) \phi_{n}(x) \tag{2}
\end{equation*}
$$

where the $\mathcal{L}$ evolution operator has the general form

$$
\begin{gather*}
\mathcal{L}(y, x)=\delta_{\sigma}[y-f(x)] \\
\delta_{\sigma}(x)=\int \delta(x-\sigma \xi) p(\xi) d \xi=(1 / \sigma) p(x / \sigma) \tag{3}
\end{gather*}
$$

For the calculations in this paper Gaussian weak noise is assumed. In the perturbative limit, $\sigma \rightarrow 0$, the evolution operator becomes

$$
\begin{equation*}
\mathcal{L}(x, y)=(1 / \sqrt{2 \pi} \sigma) e^{-[y-f(x)]^{2} / 2 \sigma^{2}} . \tag{4}
\end{equation*}
$$

The map considered here is the same as in our previous papers, a quartic map on the $(0,1)$ interval given by

$$
\begin{equation*}
f(x)=20\left[\frac{1}{16}-\left(\frac{1}{2}-x\right)^{4}\right] . \tag{5}
\end{equation*}
$$

Throughout the theory developed in previous works [4-7], the periodic orbits of the system played a major role. A periodic orbit of length $n$ was defined simply by

$$
\begin{gather*}
x_{j+1}=f\left(x_{j}\right), \quad j=1, \ldots, n  \tag{6}\\
x_{n+1}=x_{1} . \tag{7}
\end{gather*}
$$

For a repeller the leading eigenvalue of the evolution operator yields a physically measurable property of the dynamical system, the escape rate from the repeller. In the case of deterministic flows, the periodic orbit theory yields explicit formulas for the spectrum of $\mathcal{L}$ as zeros of its spectral determinant [3]. One of the most important goals of the theory related to stochastic evolution operators is to explore
the dependence of the eigenvalues $\nu$ of $\mathcal{L}$ on the noise strength parameter $\sigma$. The eigenvalues are determined by the eigenvalue condition

$$
\begin{equation*}
F[\sigma, \nu(\sigma)]=\operatorname{det}[1-\mathcal{L} / \nu(\sigma)]=0 \tag{8}
\end{equation*}
$$

where $F(\sigma, 1 / z)=\operatorname{det}(1-z \mathcal{L})$ is the spectral determinant of the evolution operator $\mathcal{L}$, which can be expressed as

$$
\begin{equation*}
\operatorname{det}(1-z \mathcal{L})=\exp \left(-\sum_{n}^{\infty} \frac{z^{n}}{n} \operatorname{Tr} \mathcal{L}^{n}\right) \tag{9}
\end{equation*}
$$

Equation (9) shows that noise dependence of the eigenvalues of the evolution operator are very closely related to the noise dependence of the trace of $\mathcal{L}^{n}$, which shall be the object of study from now on.

The trace of $\mathcal{L}^{n}$ can be expressed as

$$
\begin{equation*}
\operatorname{Tr} \mathcal{L}^{n}=\frac{1}{(\sqrt{2 \pi} \sigma)^{n}} \int d x_{1} d x_{2} \cdots d x_{n} e^{-S / \sigma^{2}} \tag{10}
\end{equation*}
$$

where

$$
\begin{gather*}
S=\frac{1}{2} \sum_{j=1}^{n}\left[x_{j+1}-f\left(x_{j}\right)\right]^{2}  \tag{11}\\
x_{n+1}=x_{1} \tag{12}
\end{gather*}
$$

In order to provide deeper insight into the forthcoming calculations, we draw a correlation between discrete Hamiltonian mechanics and our system, with the $S$ defined above playing the role of the classical action. According to Eq. (11), the least-action principle requires

$$
\begin{equation*}
x_{j}-f\left(x_{j-1}\right)-f^{\prime}\left(x_{j}\right)\left[x_{j+1}-f\left(x_{j}\right)\right]=0 \tag{13}
\end{equation*}
$$

We define

$$
\begin{equation*}
p_{j}=x_{j}-f\left(x_{j-1}\right) \tag{14}
\end{equation*}
$$

the quantity corresponding to the momentum in classical mechanics. From Eq. (13) we obtain

$$
\begin{align*}
& x_{j+1}=f\left(x_{j}\right)+p_{j+1},  \tag{15}\\
& p_{j+1}=p_{j} /\left[f^{\prime}\left(x_{j}\right)\right], \tag{16}
\end{align*}
$$

which are the equations corresponding to the classical Newtonian equations of motion. The generalized periodic orbits of length $n$ are those orbits that obey these equations and $x_{n+1}=x_{1}, p_{n+1}=p_{n}$. Those generalized periodic orbits that have nonzero momentum will control the asymptotic behavior of the corrections to $\operatorname{Tr} \mathcal{L}^{n}$ as we shall demonstrate later. The original periodic orbits defined by Eqs. (6) and (7) are those with zero momentum. The generalized periodic orbits with nonzero momentum and the original periodic orbits proliferate with growing $n$ as suggested by Fig. 1.

We introduce an integral representation of the noisy kernel, which will be of great use in the later calculations:

$$
\begin{align*}
\mathcal{L}(x, y) & =(1 / \sqrt{2 \pi} \sigma) e^{-[y-f(x)]^{2} / 2 \sigma^{2}} \\
& =\frac{1}{2 \pi} \int d k e^{-\sigma^{2} k^{2} / 2+i k[y-f(x)]} . \tag{17}
\end{align*}
$$

Using this new integral representation,


FIG. 1. The sets of original and generalized periodic orbits. Squares indicate original periodic orbits, dots indicate generalized periodic orbits, large symbols indicate orbits of length one, small symbols indicate orbits of length two.

$$
\begin{align*}
\operatorname{Tr} \mathcal{L}^{n}= & \frac{1}{(2 \pi)^{n}} \int d k^{n} d x^{n} \exp \left(-\sigma^{2} / 2 \sum_{j=1}^{n} k_{j}^{2}+i \sum_{j=1}^{n} k_{j}\left[x_{j+1}\right.\right. \\
& \left.\left.-f\left(x_{j}\right)\right]\right) \tag{18}
\end{align*}
$$

or, equivalently,

$$
\begin{align*}
\operatorname{Tr} \mathcal{L}^{n}= & \frac{1}{(2 \pi)^{n}} \int d k^{n} \int d p^{n} J(p) \exp \left(-\sigma^{2} / 2 \sum_{j=1}^{n} k_{j}^{2}\right. \\
& \left.+i \sum_{j=1}^{n} k_{j} p_{j}\right) \tag{19}
\end{align*}
$$

where $J(p)=D(x) / D(p)$ denotes the Jacobian. Since

$$
\begin{equation*}
\frac{1}{(2 \pi)^{n}} \int d k^{n} \exp \left(i \sum_{j=1}^{n} k_{j} p_{j}\right)=\prod_{j=1}^{n} \delta\left(p_{j}\right) \tag{20}
\end{equation*}
$$

we can reduce Eq. (19) to

$$
\begin{equation*}
\operatorname{Tr} \mathcal{L}^{n}=\int d p^{n} J(p) e^{\left(\sigma^{2} / 2\right) \Delta} \prod_{j=1}^{n} \delta\left(p_{j}\right)=\left.e^{\left(\sigma^{2} / 2\right) \Delta} J(p)\right|_{p_{j}=0} \tag{21}
\end{equation*}
$$

where $\Delta$ denotes the Laplacian

$$
\begin{equation*}
\Delta=\partial^{2} / \partial p_{1}^{2}+\partial^{2} / \partial p_{2}^{2}+\cdots+\partial^{2} / \partial p_{n}^{2} \tag{22}
\end{equation*}
$$

Our object of study is the Taylor expansion of Eq. (21) in the noise parameter:

$$
\begin{align*}
& \text { meter: }  \tag{23}\\
& \qquad \operatorname{Tr} \mathcal{L}^{n}=\sum_{N=0}^{\infty}\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N} \sigma^{2 N},  \tag{24}\\
& \left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N}=\left.\left(1 / 2^{N}\right)\left(\Delta^{N} / N!\right) J(p)\right|_{p_{j}=0} .
\end{align*}
$$

The $N$ th power of the Laplacian in the equation above can be

$$
\Delta^{N}=\sum_{j_{1}, \ldots, j_{n}=0}^{\infty} \frac{N!}{j_{1}!\cdots j_{n}!} \frac{\partial^{2 j_{1}}}{\partial p_{1}^{2 j_{1}}} \cdots \frac{\partial^{2 j_{n}}}{\partial p_{n}^{2 j_{n}}} \delta_{N}, \sum_{k=1}^{n} j_{k}
$$

where $\delta_{j l}$ is the Kronecker-delta. With the help of the multidimensional residue formula from complex calculus [12]

$$
\begin{align*}
\frac{\partial^{n_{1}+\cdots+n_{k}} f(z)}{\partial z_{1}^{n_{1}} \cdots \partial z_{k}^{n_{k}}}= & \frac{n_{1}!\cdots n_{k}!}{(2 \pi i)^{k}} \oint_{c_{1}} \cdots \oint_{c_{k}} \\
& \times \frac{f(\xi) d \xi_{1} \cdots d \xi_{k}}{\left(\xi_{1}-z_{1}\right)^{n_{1}+1} \cdots\left(\xi_{k}-z_{k}\right)^{n_{k}+1}} \tag{26}
\end{align*}
$$

we obtain

$$
\begin{align*}
\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N}= & \frac{1}{(2 \pi i)^{n} 2^{N}} \sum_{j_{1}, \ldots, j_{n}=0}^{\infty} \frac{\left(2 j_{1}\right)!}{j_{1}!} \cdots \frac{\left(2 j_{n}\right)!}{j_{n}!} \\
& \times \delta_{N,} \sum_{k=1}^{n} j_{k} \oint_{C_{1}} \ldots \oint_{C_{n}} \frac{J(p) d p_{1} \cdots d p_{n}}{p_{1}^{2 j_{1}+1} \cdots p_{n}^{2 j_{n}+1}} . \tag{27}
\end{align*}
$$

The contours are around the $p_{j}=0$ points. The integrals can be transformed back to contour integrals in the original $x_{j}$ variables, and the contours will be placed around the original periodic orbits of the system defined by Eqs. (6) and (7), since it is these orbits that fulfil the $p_{j}=0$ conditions. From now on we shall restrict our calculations to the asymptotic large $N$ limit. We will replace the summations in Eq. (27) by integrals and then use the saddle-point method to get a compact formula for $\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N}$. We approximate the factorials via the Stirling formula [13] as

$$
\begin{align*}
\frac{\left(2 j_{k}\right)!}{j_{k}!} & \simeq \frac{\left.\left(2 j_{k} / e\right)\right)^{2 j_{k}} \sqrt{4 \pi j_{k}}}{\left(j_{k} / e\right)^{j_{k}} \sqrt{2 \pi j_{k}}} \\
& =2^{2 j_{k}+1 / 2} j_{k}^{j_{k}} e^{-j_{k}}=2^{1 / 2} \exp \left[2(\ln 2) j_{k}+j_{k} \ln j_{k}-j_{k}\right] \tag{28}
\end{align*}
$$

Using Eq. (28) and an integral representation of the delta function, we get

$$
\begin{align*}
\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N} \simeq & 2^{(n / 2)-N} /\left[(2 \pi i)^{n} 2 \pi\right] \\
& \times \sum_{j_{1}, \ldots, j_{n}=0}^{\infty} \int d t \oint_{c_{1}} \ldots \oint_{c_{n}} d x_{1} \cdots d x_{n} \\
& \times \exp \left[i t\left(N-\sum_{k=1}^{n} j_{k}\right)+(2 \ln 2-1) \sum_{k=1}^{n} j_{k}\right. \\
& \left.+\sum_{k=1}^{n} j_{k} \ln j_{k}+\sum_{k=1}^{n} \ln \left[x_{k}-f\left(x_{k-1}\right)\right]\left(2 j_{k}+1\right)\right] . \tag{29}
\end{align*}
$$

Now we replace $j_{k}$ with the new variables $y_{k}=j_{k} / N$ and in the asymptotic ( $N$ large) limit approximate the summations for $y_{k}$ by integrals for $y_{k}$ as

$$
\begin{align*}
\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N} \simeq & \frac{2^{n / 2-N} N^{n}}{(2 \pi i)^{n} 2 \pi} \int_{0}^{\infty} d y_{1} \cdots \int_{0}^{\infty} d y_{n} \\
& \times \int d t \oint_{c_{1}} \ldots \oint_{c_{n}} d x_{1} \cdots d x_{n} \\
& \times \exp \left[i t\left(N-N \sum_{k=1}^{n} y_{k}\right)+N(2 \ln 2-1) \sum_{k=1}^{n} y_{k}\right. \\
& +N \sum_{k=1}^{n} y_{k} \ln \left(N y_{k}\right)+\sum_{k=1}^{n} \ln \left[x_{k}-f\left(x_{k-1}\right)\right] \\
& \left.\times\left(2 N y_{k}+1\right)\right] . \tag{30}
\end{align*}
$$

We evaluate the $y$ integrals with the saddle point method to get

$$
\begin{align*}
\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N} \simeq & \frac{2^{-N+n / 2}}{(2 \pi)^{n / 2} i^{n} 2 \pi} \int d t \oint_{c_{1}} \ldots \oint_{c_{n}} d x_{1} \cdots d x_{n} \\
& \times \exp \left[i t\left(N+\frac{n}{2}\right)-e^{i t} \frac{S}{2}\right] . \tag{31}
\end{align*}
$$

Next we apply the saddle point method to the integral in $t$ as well, asymptotically resulting in

$$
\begin{align*}
\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N} \simeq & \frac{N^{(n-1) / 2}}{2^{2 N+1 / 2}(2 \pi)^{(n+1) / 2} i^{n+1}} \frac{(2 N)!}{N!} \\
& \times \int d x^{n} e^{-(N+n / 2) \ln (S)} \tag{32}
\end{align*}
$$

The last step is to evaluate the contour integrals in the $x_{k}$ variables. We deform the contours until the saddle points are reached so the contours run along the routes of steepest descent. The leading contribution comes from those saddle points that fulfil the following equation

$$
\begin{equation*}
(1 / S)\left\{x_{j}^{*}-f\left(x_{j-1}^{*}\right)-\left[x_{j+1}^{*}-f\left(x_{j}^{*}\right)\right] f^{\prime}\left(x_{j}^{*}\right)\right\}=0 . \tag{33}
\end{equation*}
$$

By comparing Eqs. (33) and (13) one can see that the saddle points are all generalized periodic orbits of the system. Since the contours ran originally around the orbits with zero momentum, these do not come into play as saddle points. The second derivative matrix is

$$
\begin{equation*}
-[N+(n / 2)](1 / S) D^{2} S, \tag{34}
\end{equation*}
$$

where $D^{2} S$ denotes the second derivative matrix of $S$

$$
\begin{equation*}
\left(D^{2} S\right)_{i j}=\left(\partial^{2} S\right) /\left(\partial x_{i} \partial x_{j}\right) \tag{35}
\end{equation*}
$$

This would be the matrix to deal with if we had taken the saddle-point approximation of Eq. (10) directly. We reorganize the prefactor in Eq. (32) with the use of the Stirling formula [13] and the result of the saddle-point (sp) integration is written as
$\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N} \simeq \sum_{s p} \frac{N^{(n-1) / 2}}{2 \pi i} \frac{\Gamma\left(N+\frac{1}{2}\right)}{[N+(n / 2)]^{n / 2}} \frac{S_{p}^{-N}}{\sqrt{\operatorname{det} D^{2} S_{p}}}$,


FIG. 2. The ratio of $\left(\operatorname{Tr} \mathcal{L}^{2}\right)_{N}$ calculated via the asymptotic formula (36) to its value computed by numerical integration.
which is our main result. For $n=1$ this formula gives the result of [7] as it should.

Finally, we draw attention to the close connection between the generalized periodic orbits of the system and $D^{2} S$. The stability matrix of a general periodic orbit (po) is expressed as

$$
\begin{align*}
& J=J_{1} \cdot J_{2} \cdot J_{3} \cdots J_{n},  \tag{37}\\
& J_{k}=\left(\begin{array}{cc}
f^{\prime}\left(x_{k}\right)-\frac{p_{k}}{\left[f^{\prime}\left(x_{k}\right)\right]^{2}} f^{\prime \prime}\left(x_{k}\right) & \frac{1}{f^{\prime}\left(x_{k}\right)} \\
-\frac{p_{k}}{\left[f^{\prime}\left(x_{k}\right)\right]^{2}} f^{\prime \prime}\left(x_{k}\right) & \frac{1}{f^{\prime}\left(x_{k}\right)}
\end{array}\right) . \tag{38}
\end{align*}
$$

The determinant of $D^{2} S$ can be expressed with the help of the stability matrix as

$$
\begin{equation*}
\operatorname{det} D^{2} S_{p}=\operatorname{det}\left(J_{p}-1\right) \tag{39}
\end{equation*}
$$

This way we reformulate Eq. (36) as

$$
\begin{equation*}
\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N} \simeq \frac{N^{(n-1) / 2}}{2 \pi} \frac{\Gamma\left(N+\frac{1}{2}\right)}{[N+(n / 2)]^{n / 2}} \sum_{p o} \frac{e^{-N \ln S_{p}}}{\sqrt{\operatorname{det}\left(1-J_{p}\right)}} \tag{40}
\end{equation*}
$$

where the summation runs over generalized periodic orbits, with nonzero momentum. This is fully analogous to a trace formula and is our main result.

Finally we turn toward testing our result obtained so far. In [7] we developed a contour integral method to calculate high-order noise corrections to the trace of $\mathcal{L}$. We showed that the agreement between the exact results and a formula that coincides with the Eq. (36) in the $n=1$ case is very good. Now we step ahead and produce numerically highorder noise corrections to the trace of $\mathcal{L}^{2}$. We shall start from Eq. (27) by transforming the integrals in $p$ back to integrals in $x$ as

$$
\begin{align*}
\left(\operatorname{Tr} \mathcal{L}^{n}\right)_{N}= & \frac{1}{(2 \pi i)^{n} 2^{N}} \sum_{j_{1}, \ldots, j_{n}=0}^{\infty} \frac{\left(2 j_{1}\right)!}{j_{1}!} \cdots \frac{\left(2 j_{n}\right)!}{j_{n}!} \\
& \times \delta_{N, \Sigma_{k=1}^{n} j_{k}} \oint_{c_{1}} \cdots \oint_{c_{n}} \\
& \times \frac{d x_{1} \cdots d x_{n}}{\left[x_{1}-f\left(x_{n}\right)\right]^{2 j_{1}+1} \cdots\left[x_{n}-f\left(x_{n-1}\right)\right]^{2 j_{n}+1}} . \tag{41}
\end{align*}
$$

The contours at Eq. (27) were around the $p_{j}=0$ points, so the contours above are placed around the original periodic orbits of the system, defined by Eqs. (6) and (7). These contour integrals can be evaluated numerically. Figure 2 shows the ratio of $\left(\operatorname{Tr} \mathcal{L}^{2}\right)_{N}$ obtained from Eq. (36) and evaluated via the procedure described above as a function of $N$.

In summary, we have studied the evolution operator for a discrete Langevin equation with a strongly hyperbolic classical dynamics and a Gaussian noise distribution. Using an integral representation of the evolution operator $\mathcal{L}$ we have revealed the asymptotic behavior of the corrections to the trace of $\mathcal{L}^{n}$. This behavior is governed by subdominant terms corresponding to terms previously neglected in the perturbative expansion, and a fully analogous trace formula can be derived for the late terms in the noise extension series of the trace of $\mathcal{L}^{n}$.
G.V. and G.P. gratefully acknowledge the financial support of the Hungarian Ministry of Education, OMFB, OTKA Grant No. T25866. G.P., G.V., and A.V. were also partially supported by the French Ministère des Affaires Étrangères.
[1] E. Aurell et al., Nonlinearity 3, 325 (1990).
[2] E. Aurell et al., Nonlinearity 3, 361 (1990).
[3] P. Cvitanović et al., Classical and Quantum Chaos (Niels Bohr Institute, Copenhagen, 1999), http://www.nbi.dk/ ChaosBook/.
[4] P. Cvitanović et al., J. Stat. Phys. 93, 981 (1998).
[5] P. Cvitanović, C.P. Dettmann, R. Mainieri, and G. Vattay, Nonlinearity 12, 939 (1998).
[6] P. Cvitanović et al., Phys. Rev. E 60, 3936 (1999).
[7] N. Søndergaard et al., J. Stat. Phys. 101, 385 (2000).
[8] D. Alonso and P. Gaspard, Chaos 3, 601 (1993); P. Gaspard and D. Alonso, Phys. Rev. A 47, R3468 (1993).
[9] G. Vattay and P.E. Rosenqvist, Phys. Rev. Lett. 76, 335 (1996).
[10] G. Vattay, Phys. Rev. Lett. 76, 1059 (1996).
[11] M.C. Gutzwiller, Chaos in Classical and Quantum Mechanics (Springer, New York, 1990).
[12] S. Bochner and W.T. Martin, Several Complex Variables (Princeton University Press, Princeton, NJ, 1948), Chap. 2, p. 33, Formula 14.
[13] M. Abramowitz and I. A. Stegun, Handbook of Mathematical Functions with Formulas, Graphs and Mathematical Tables (Dover, New York, 1972), Chap. 6, p. 257, Formula 6.1.37.

